

The LIEDR model - recent and future improvements

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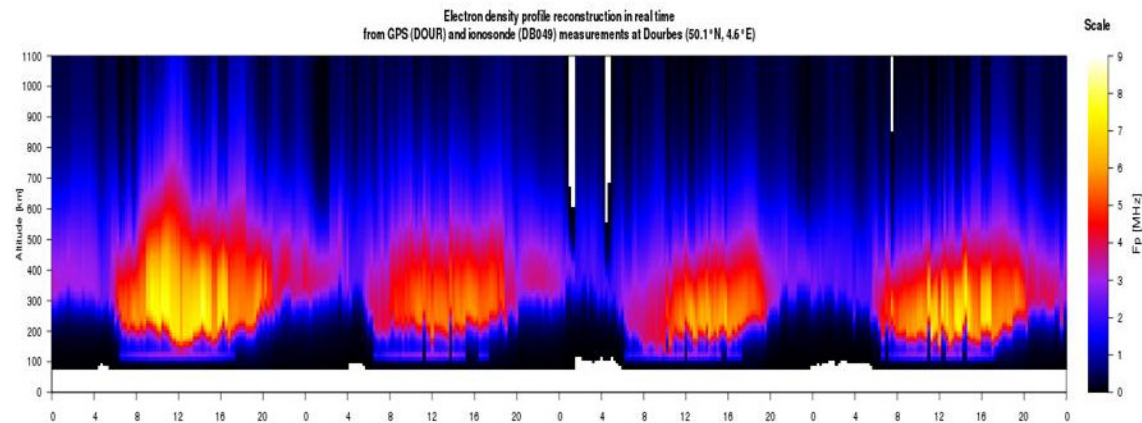
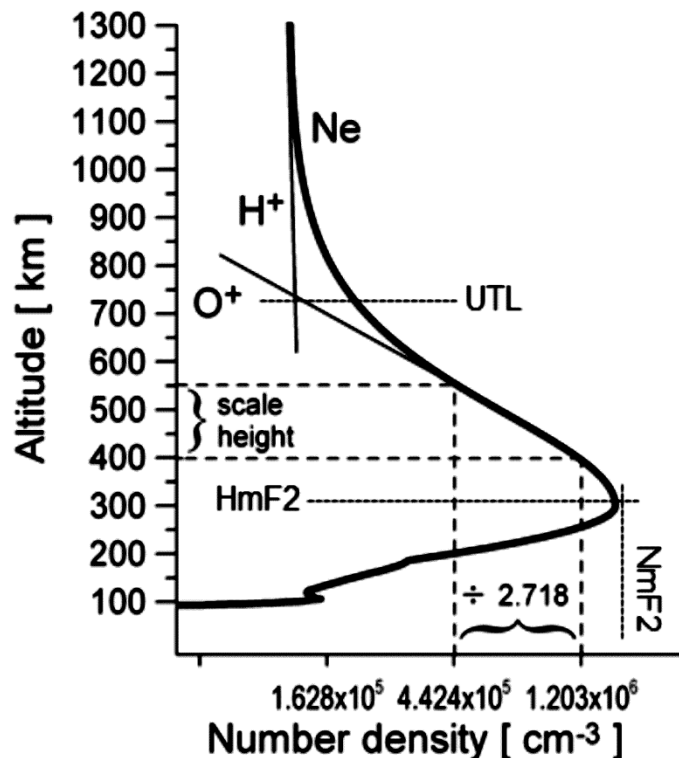
Solar-Terrestrial
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Overview

- Introduction
- The topside profiler problem
- The available topside data
- Correlations with physical drivers
- Correlations with measured ionospheric parameters
- Further developments

Introduction

The goal of LIEDR is to reconstruct the electron density profile for the ionosphere:



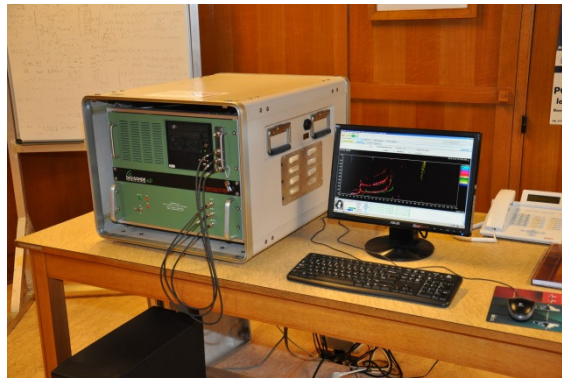
LIEDR reconstructed electron density profile for four days, showing several failed reconstructions

N_e is reconstructed from 80km up to 1100km.

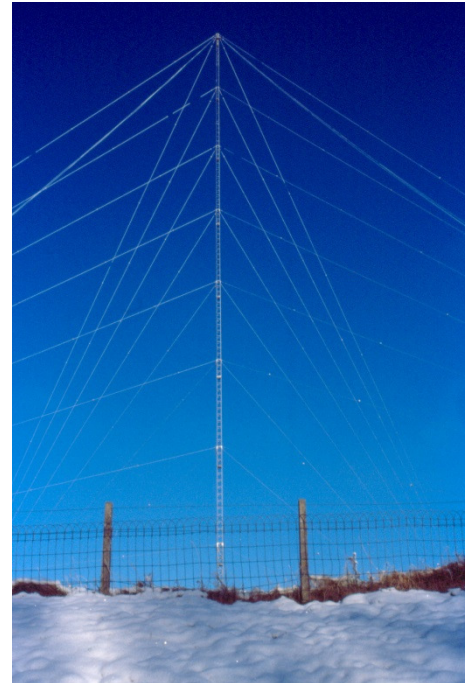
LIEDR inputs

- 1) Parameters from ground based ionosonde
- 2) Total (vertical) electron content, obtained from TEC maps (provided by ROB)
- 3) Solar activity index (F10.7) and geomagnetic measurements
- 4) An empirical model for the transition height (UTL)

The ionosonde



Sounder,
transmission
antenna and
receiver antenna



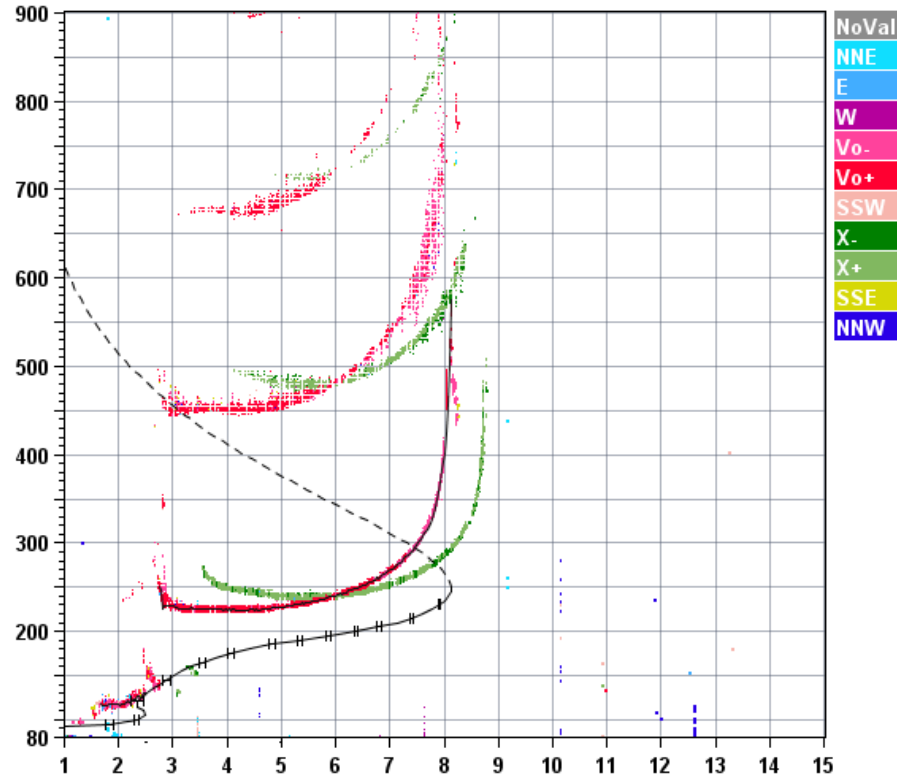
With the new digisonde, installed in 2011, these measurements are now as good as they can get.

The ionosonde



Station YYYY DAY DDD HHMMSS P1 FFS S AXN PPS IGA PS
 Dourbes 2012 Jan07 007 094500 RSF 005 2 513 100 03+ B1

foF2	8.128
foF1	N/A
foF1p	N/A
foE	2.50
foEp	2.38
fxI	8.84
foEs	2.78
fmin	1.66
<hr/>	
MUF(D)	27.52
M(D)	3.39
D	N/A
<hr/>	
h'F	223.0
h'F2	223.0
h'E	115.0
h'Es	134.8
<hr/>	
hmF2	248.8
hmF1	N/A
hmE	105.5
yF2	78.8
yF1	N/A
yE	15.2
B0	61.3
B1	3.45
<hr/>	
C-level	11
<hr/>	
Auto:	
Artist5	
500200	



D 100 200 400 600 800 1000 1500 3000 [km]
 MUF 8.8 8.9 9.3 10.0 11.0 12.4 16.6 27.5 [MHz]
 DB049_2012007094500.RSF / 437fx512h 32 kHz 2.5 km / DPS-4D DB049 049 / 50.1 N 4.6 E Ion2Png v. 1.3.17

A typical echo obtained by one frequency sweep of the ionosonde.

Black line is the calculated electron density profile.

Evaluation of the software doing this calculation still needs to be done.

LIEDR inputs (again)

The main problem:

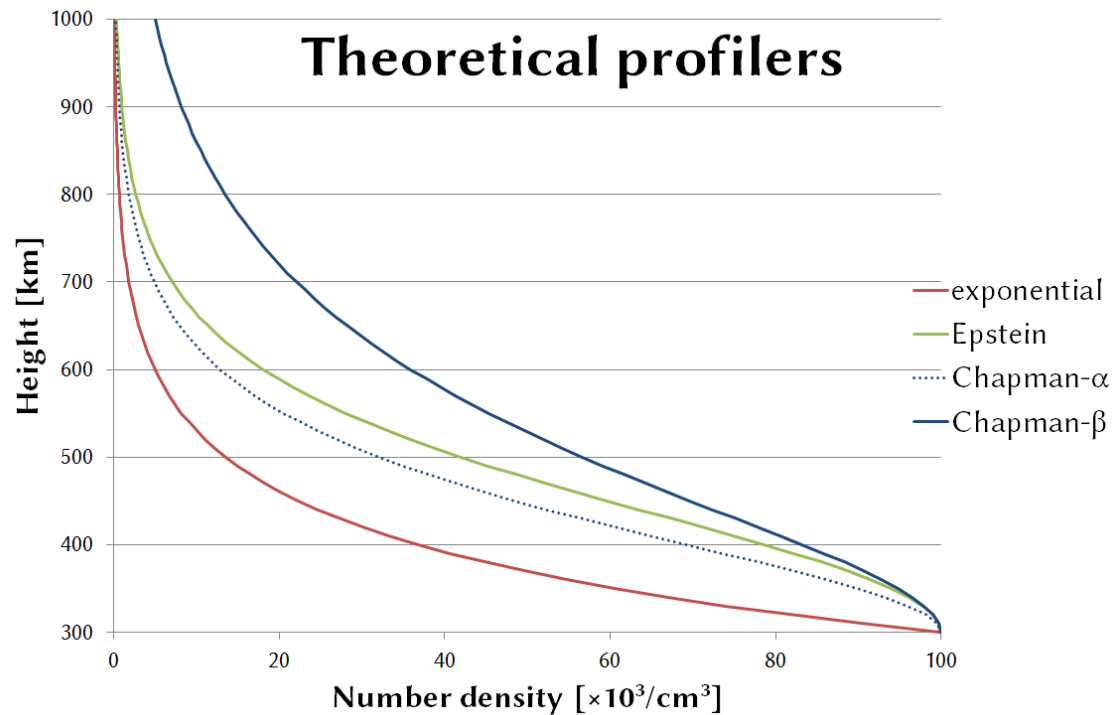
Ground based ionosonde can only measure detailed profiles up to the F_2 peak, for the topside only total electron content can be measured.

Question: how are these electrons distributed between F_2 height and UTL?

Additional input needed: a model for the topside distribution

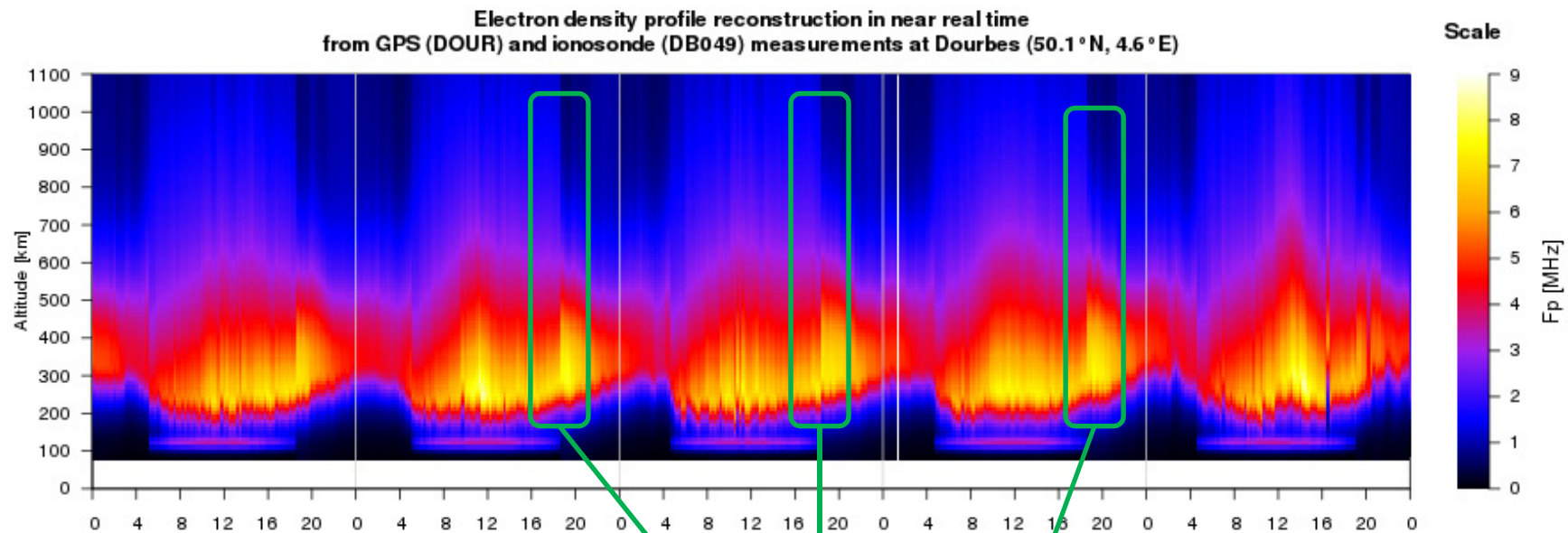
Topside profilers

Different topside profilers are commonly used: exponential, Epstein and Chapman- α and - β .



Day/night selection of profilers

Different profilers are used during day and night, however the transition is problematic.



Obviously unphysical discontinuities occur at day-night transition.

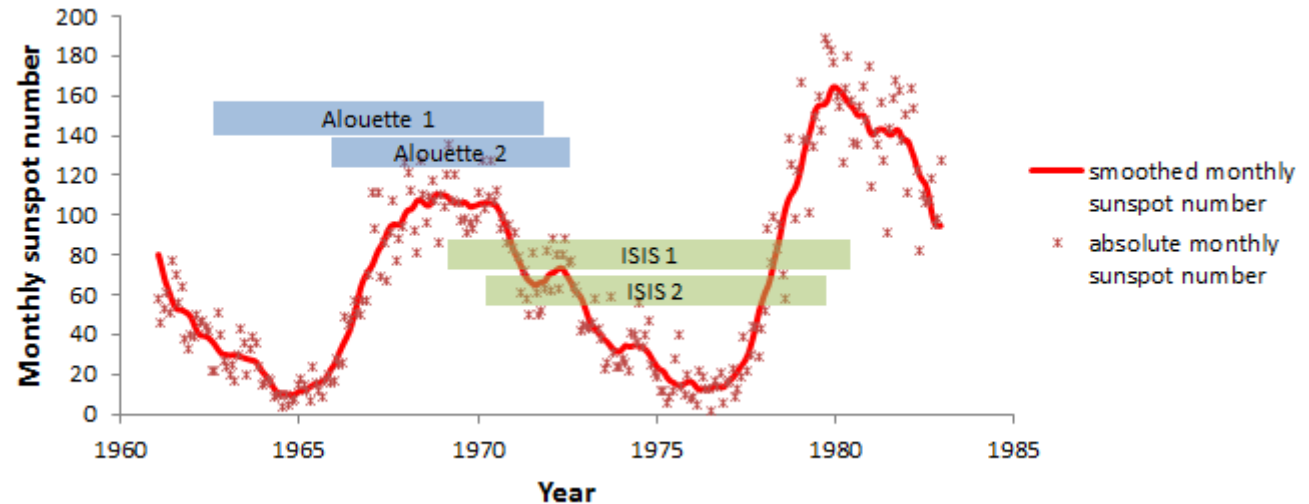
⇒ Better criteria for profiler selection are important

Topside ionosonde data

- We need to further investigate which profiler is best suited for use in what conditions.
- Relevant factors include time of day, but also season, solar activity, magnetic activity and magnetic coordinates.
- Data used is historical database, provided by NSSDC, with measurements from four satellites: Alouette 1 & 2 and ISIS 1 & 2.
- 170000+ profiles can each be fitted with the four profilers.

Data coverage & problems

The data covers all latitudes and longitudes, and spans over more than a complete solar cycle.



Problems:

- 1) The data is not uniformly distributed in either time or space.
- 2) There are some profiles that are obviously erroneous and have to be excluded from analyses.

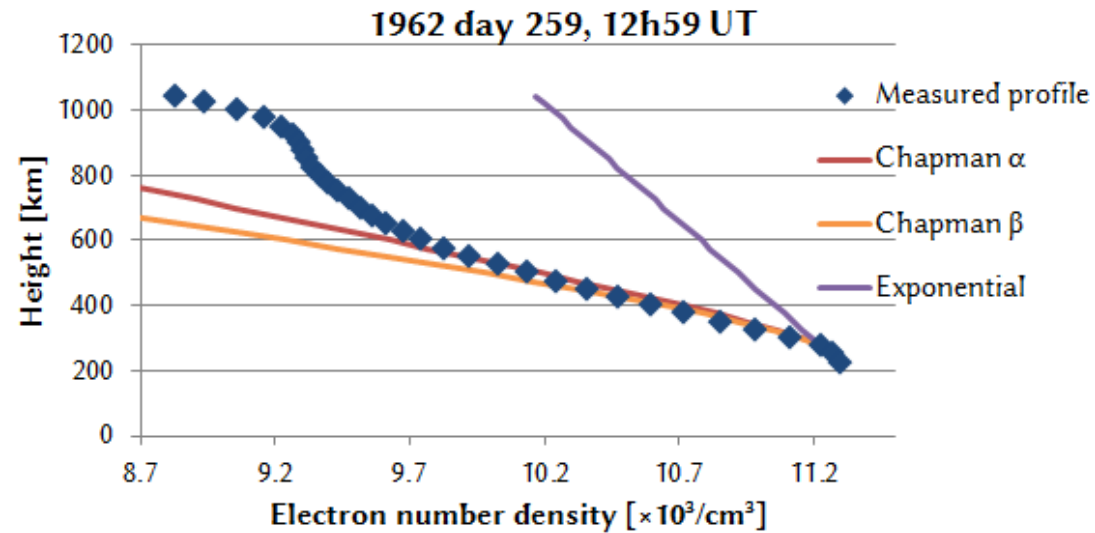
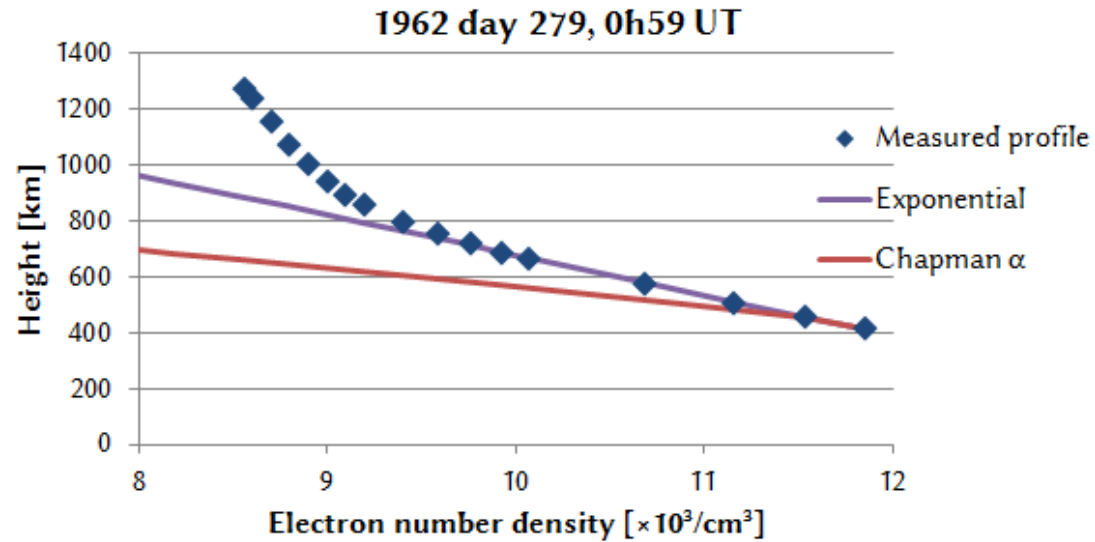
Data coverage & problems

- 3) To ensure no erroneous data is included, we select profiles based on the proximity of the peak to the IRI prediction. In disturbed conditions, this can exclude correct data, too.
- 4) We only want profiles that cover the whole topside – from $h_m F_2$ to UTL. This is the main reason for profiles being excluded from our study.

This last criterion also causes an important bias: profiles with higher UTL are more likely to not go high enough, due to the UTL being above the satellite.

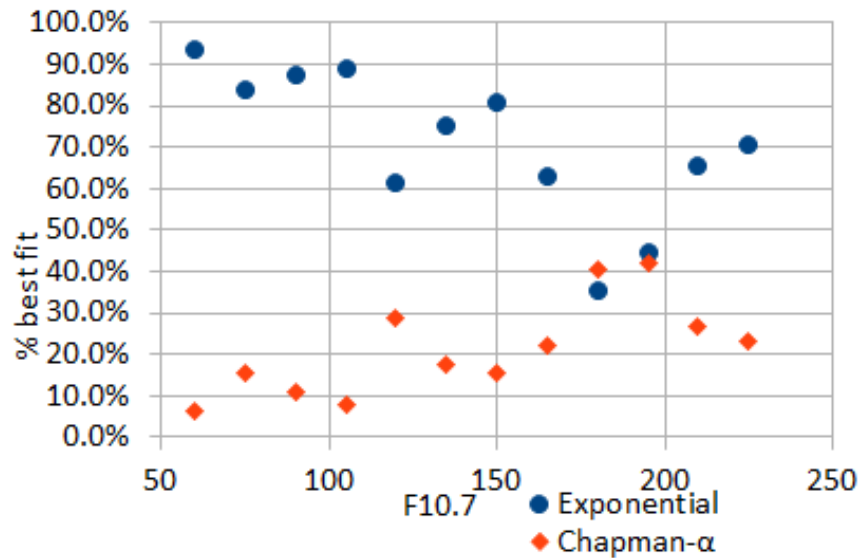
Examples

One profile best fitted by an exponential curve (top), the other by a Chapman profile (bottom)

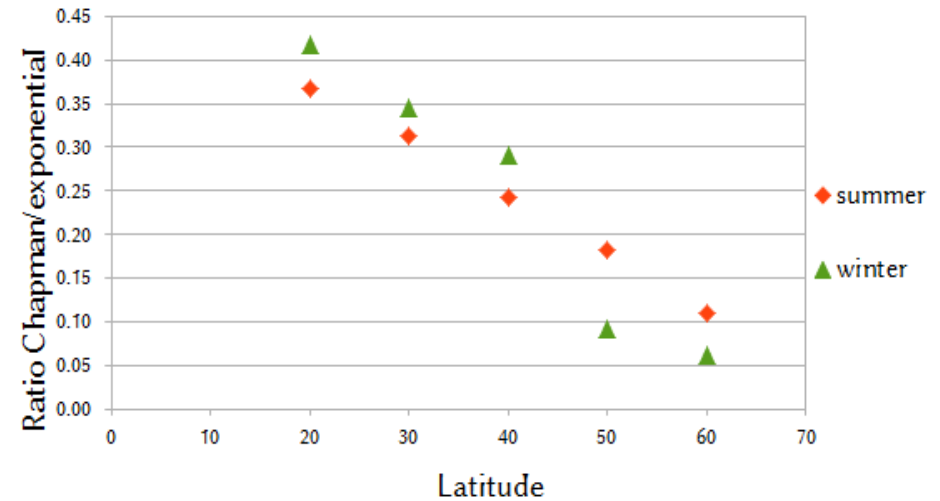


Effects of physical drivers

Effect of F10.7 (daytime)



Effect of latitude (and season)



Effects of physical drivers

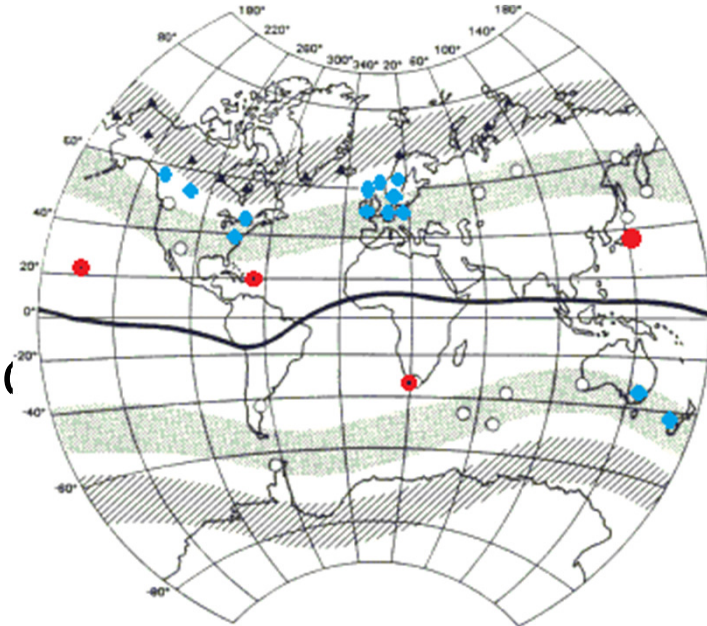
Time of day

	0-2	2-4	4-6	6-8	8-10	10-12	12-14	14-16	16-18	18-20	20-22	22-24	
80°-90°	0.13			0.56	0.56	1.38	1.09	5.50	0.24	0.80	0.10	0.00	80°-90°
70°-80°				0.31	0.21	0.36	0.68	0.23	0.09	0.08		0.20	70°-80°
60°-70°	0.21	0.26		0.43	0.16	0.25	0.37	0.29	0.59	0.61	0.39	0.15	60°-70°
50°-60°	0.16	0.22	0.31	0.24	0.29	0.33	0.28	0.15	0.38	0.50	0.23	0.41	50°-60°
40°-50°	0.17	0.16	0.24	0.32	0.17	0.24	0.20	0.10	0.21	0.22	0.37	0.33	40°-50°
30°-40°	0.11	0.13	0.21	0.33	0.17	0.31	0.28	0.13	0.21	0.27	0.39	0.43	30°-40°
20°-30°	0.29	0.29	0.29	0.35	0.25	0.50	0.43	0.16	0.25	0.28	0.47	0.44	20°-30°
10°-20°	0.53	0.74	0.57	0.48	0.50	0.48	0.45	0.30	1.02	0.76	1.01	0.57	10°-20°
0°-10°	0.87	0.91	0.61	0.94	1.26	1.57	0.98	1.31	1.33	2.05	1.08	0.73	0°-10°
80°-90°	0.24	0.38			0.73	1.09	4.00	20.00					80°-90°
70°-80°	0.24	0.54			0.27	0.27	1.40	0.96	1.43				70°-80°
60°-70°	0.33	0.59		0.38		0.50	0.89	0.41					60°-70°
50°-60°	0.36	0.42	0.20	0.51	0.62	0.69	0.62	0.52	0.34	0.55	0.29	0.61	50°-60°
40°-50°	0.41	0.42	0.14	0.45	0.41	0.73	0.56	0.41	0.36	0.53	0.36	0.53	40°-50°
30°-40°	0.33	0.41	0.12	0.31	0.40	0.25	0.51	0.43	0.35	0.42	0.57	0.41	30°-40°
20°-30°	0.29	0.25	0.20	0.28	0.52	0.38	0.80	0.72	0.64	0.41	0.57	0.43	20°-30°
10°-20°	0.52	0.78	0.76	0.77	1.09	0.73	1.05	0.92	0.64	0.48	0.53	1.08	10°-20°
0°-10°	0.94	0.91	1.41	1.59	0.84	1.23	0.93	0.43	0.27	1.00	1.45	1.00	0°-10°

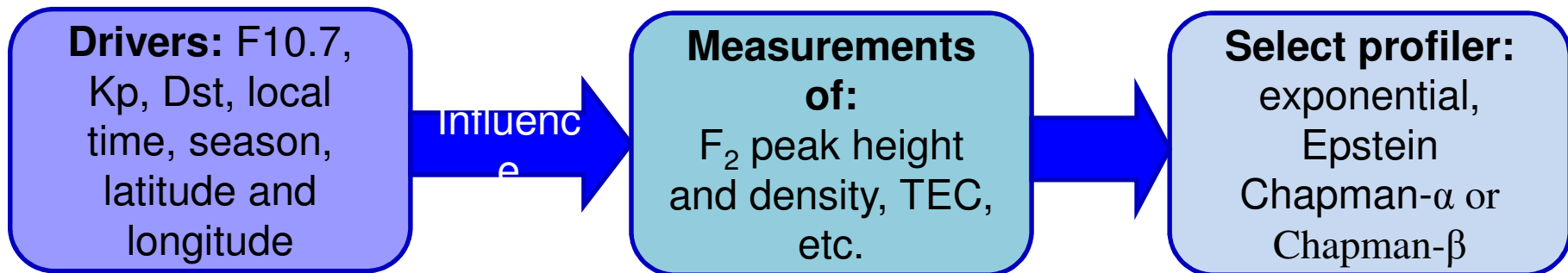
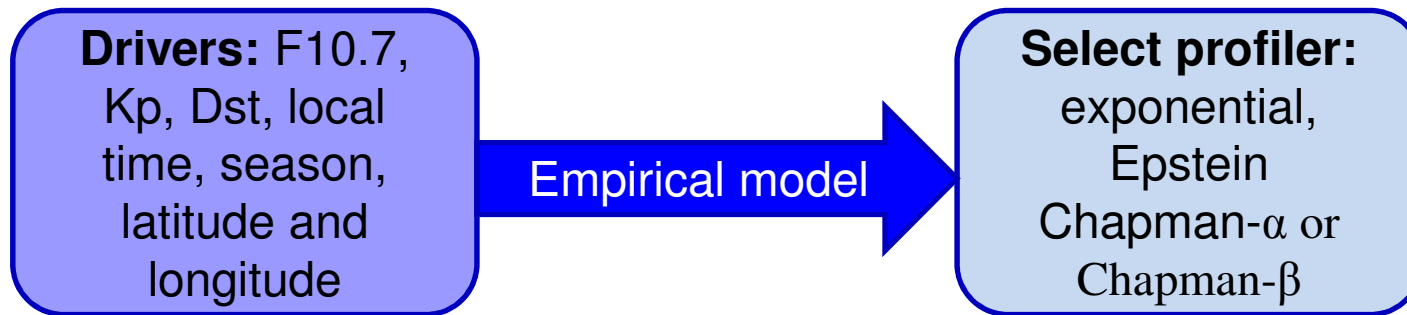
Ratio of non-exponential to exponential profiles, for winter (top) and summer (bottom), by magnetic latitude and local time

The problem with empirical models

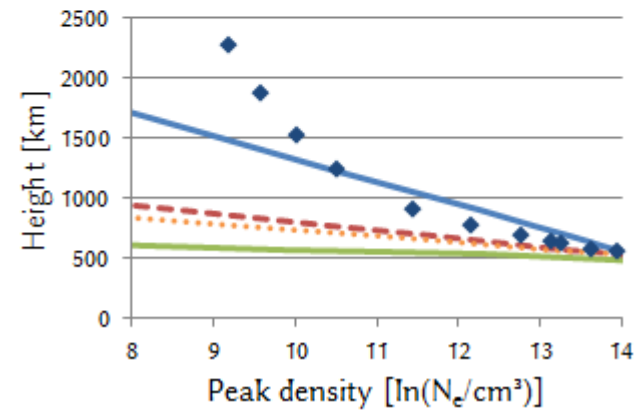
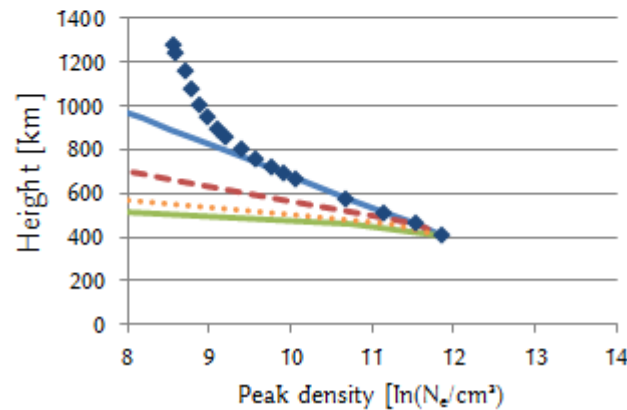
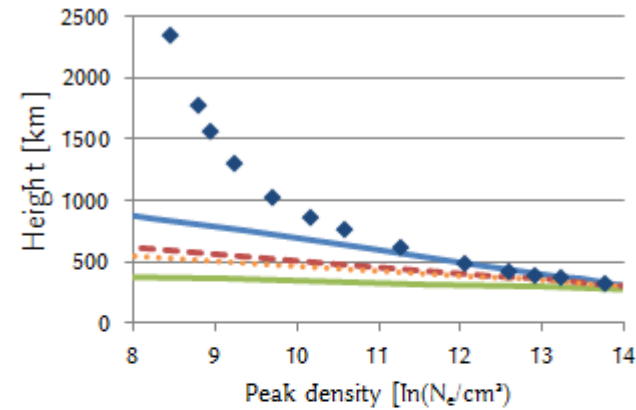
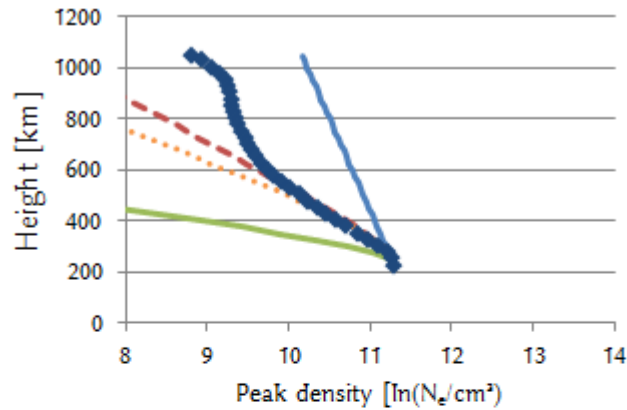
- Effects of the physical drivers are visible, but not usable for the selection of a topside profiler.
- Reasons: correlation between drivers, biases, and intrinsic problems with solar and magnetic indices.
- Better results might be obtained by using measured parameters instead of physical drivers.



The solution: measurement based

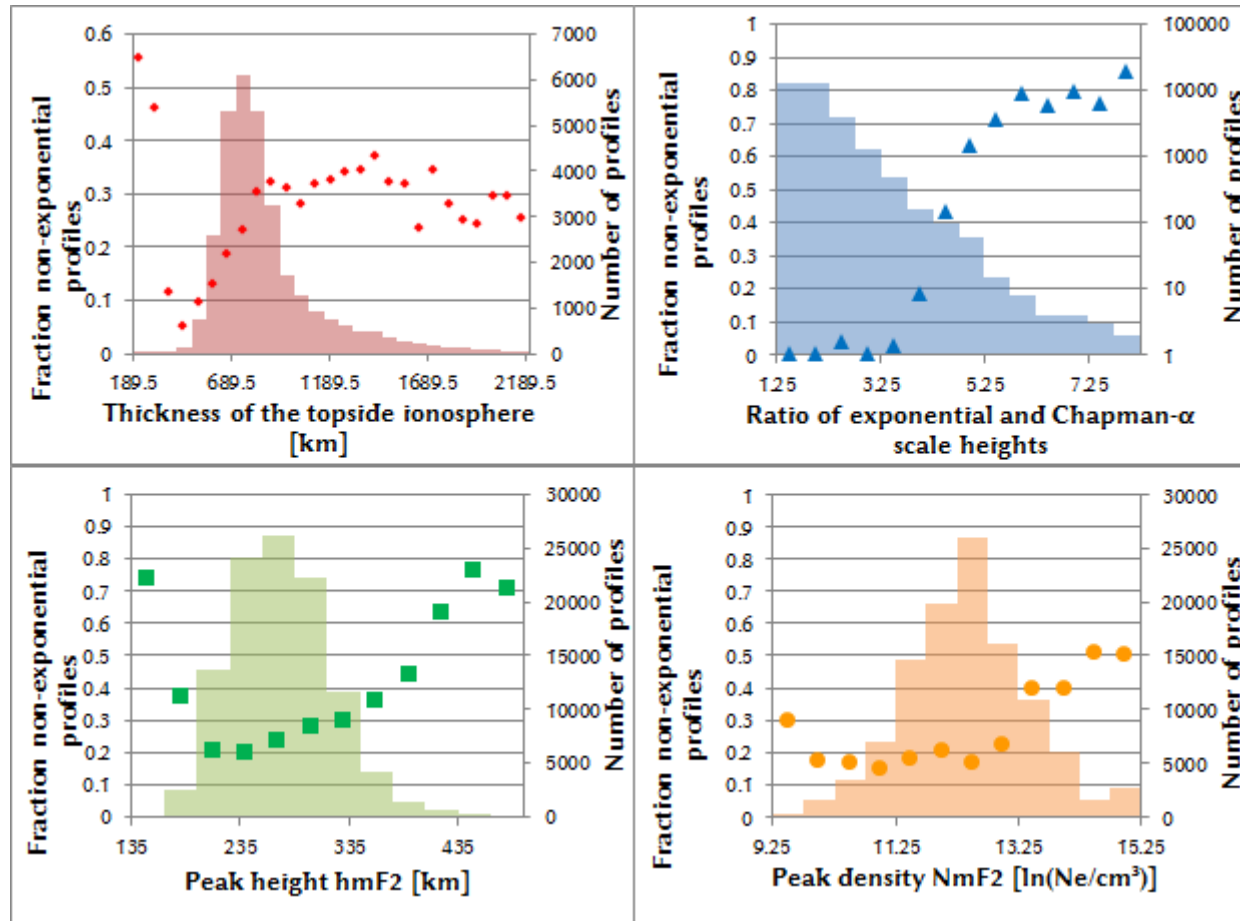


Examples

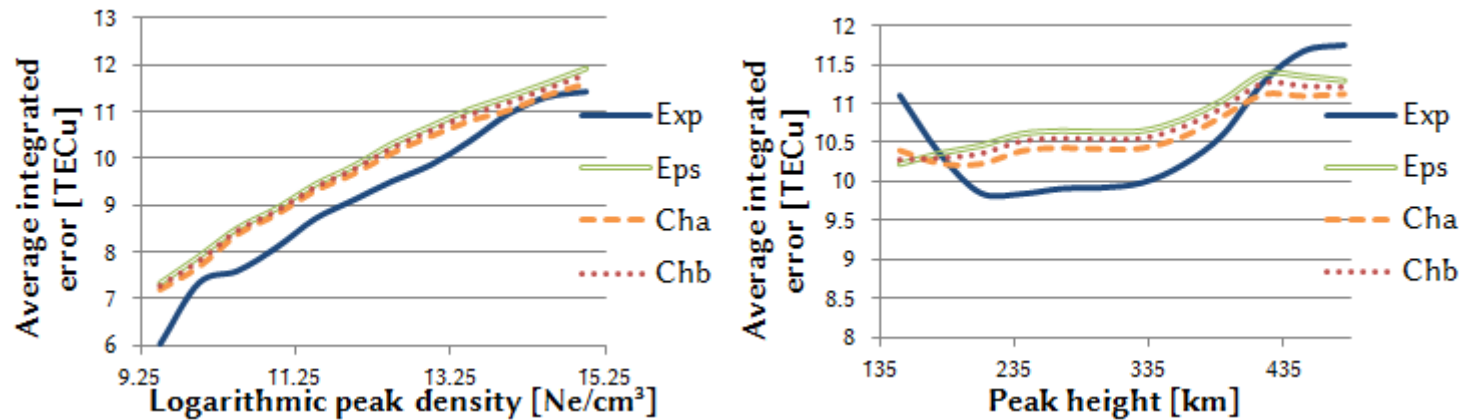


Profiles with extremely large or small values for $h_m F_2$ and $N_m F_2$ are fitted best by different profilers

Relation to measured parameters



Average integrated error



It seems that mainly the height of the F₂ peak is useful for the selection of a topside profiler

Future developments

- Implement the selection of topside profiler based of peak characteristics (and possibly TEC)
- If feasible, implement compound profiles or variable scale heights
- Evaluate the new scaling software for the ionosonde data
- Use real time, local GPS data instead of TEC maps
- At a later stage, review the model for the UTL